Supported in part by Google UK, Ltd



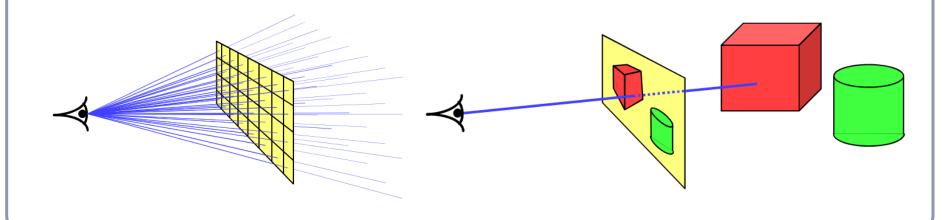
"Cornell Box" by Steven Parker, University of Utah.

A tera-ray monte-carlo rendering of the Cornell Box, generated in 2 CPU years on an Origin 2000. The full image contains 2048 x 2048 pixels with over 100,000 primary rays per pixel (317 x 317 jittered samples). Over one trillion rays were traced in the generation of this image

## Ray tracing

- A powerful alternative to polygon scan-conversion techniques
- An elegantly simple algorithm:

Given a set of 3D objects, shoot a ray from the eye through the center of every pixel and see what it hits.



### The algorithm

Select an eye point and a screen plane.

for (every pixel in the screen plane):

Find the ray from the eye through the pixel's center.

for (each object in the scene):

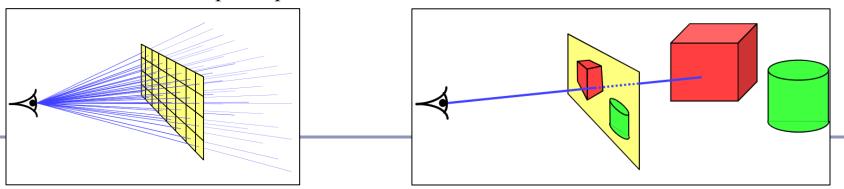
if (the ray hits the object):

if (the intersection is the nearest (so far) to the eye):

Record the intersection point.

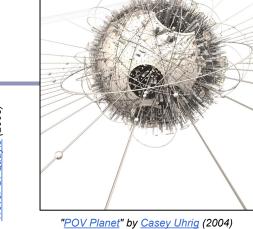
Record the color of the object at that point.

Set the screen plane pixel to the nearest recorded color.

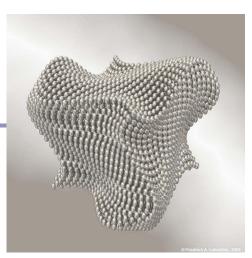


## Examples









"Dancing Cube" by Friedrich A. Lohmueller (2003)







#### It doesn't take much code

# The basic algorithm is straightforward, but there's much room for subtlety

- Refraction
- Reflection
- Shadows
- Anti-aliasing
- Blurred edges
- Depth-of-field effects
- ...



typedef struct{double x,y,z;}vec;vec U,black,amb={.02,.02,.02}; struct sphere{vec cen,color; double rad,kd,ks,kt,kl,ir;}\*s,\*best  $sph[]=\{0.,6.,.5,1.,1.,1.,.9,.05,.2,.85,0.,1.7,-1.,8.,-.5,1.,.5\}$ ,.2,1.,.7,.3,0.,.05,1.2,1.,8.,-.5,.1,.8,.8,1.,.3,.7,0.,0.,1.2,3 .,-6.,15.,1.,.8,1.,7.,0.,0.,0.,.6,1.5,-3.,-3.,12.,.8,1.,1.,5.,0 .,0.,0.,.5,1.5,};int yx;double u,b,tmin,sqrt(),tan();double vdot(vec A, vec B) {return A.x\*B.x+A.y\*B.y+A.z\*B.z;}vec vcomb( double a, vec A, vec B)  $\{B.x+=a*A.x;B.y+=a*A.y;B.z+=a*A.z;return\}$ B; \text{vec vunit(vec A) \text{return vcomb(1./sqrt(vdot(A,A)),A,black);}} struct sphere\*intersect(vec P, vec D) {best=0;tmin=10000;s=sph+5; while (s-->sph) b=vdot (D, U=vcomb(-1., P, s->cen)), u=b\*b-vdot(U, U) +s->rad\*s->rad,u=u>0?sqrt(u):10000,u=b-u>0.000001?b-u:b+u,tmin= u>0.00001&&u<tmin?best=s,u:tmin;return best;}vec trace(int level, vec P, vec D) {double d, eta, e; vec N, color; struct sphere\*s, \*1; if(!level--) return black; if(s=intersect(P,D)); else return amb; color=amb; eta=s->ir; d=-vdot(D, N=vunit(vcomb(-1., P=vcomb(  $tmin, D, P), s\rightarrow cen))); if (d<0) N=vcomb(-1., N, black), eta=1/eta, d=$ -d; l=sph+5; while (l-->sph) if ((e=l->kl\*vdot(N, U=vunit(vcomb(-1.,P ,1->cen))))>0&&intersect(P,U)==1)color=vcomb(e,1->color,color); U=s->color;color.x\*=U.x;color.y\*=U.y;color.z\*=U.z;e=1-eta\*eta\*( 1-d\*d); return vcomb(s->kt,e>0?trace(level,P,vcomb(eta,D,vcomb( eta\*d-sqrt(e),N,black))):black,vcomb(s->ks,trace(level,P,vcomb( 2\*d, N, D)), vcomb(s->kd, color, vcomb(s->kl, U, black))));}main(){int d=512; printf("%d %d\n",d,d); while(yx<d\*d){U.x=yx%d-d/2;U.z=d/2yx++/d;U.y=d/2/tan(25/114.5915590261);U=vcomb(255.,trace(3,black, vunit(U)), black); printf("%0.f %0.f %0.f\n", U.x, U.y, U.z);} }/\*minray!\*/

Paul Heckbert's 'minray' ray tracer, which fit on the back of his business card. (circa 1983)

#### Running time

The ray tracing time for a scene is a function of

(num rays cast) x
(num lights) x
(num objects in scene) x
(num reflective surfaces) x
(num transparent surfaces) x
(num shadow rays) x
(ray reflection depth) x ...



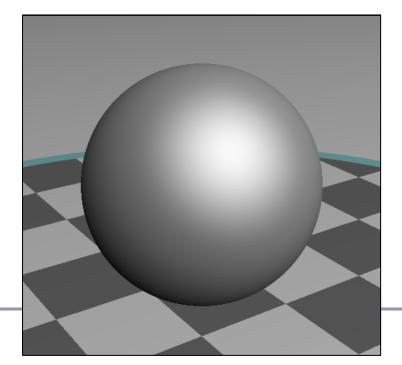
Image by nVidia

Contrast this to polygon rasterization: time is a function of the number of elements in the scene times the number of lights.

#### Recall: illumination

The *total illumination at P* is:

$$I(P) = k_A + k_D (N \cdot L) + k_S (R \cdot E)^n$$
  
summed over all lights  $L$ .

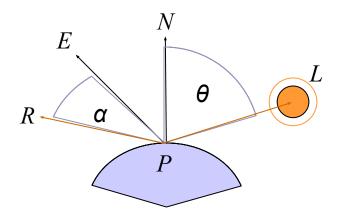


Ambient light:  $k_A$ 

Diffuse light:  $k_D(N \cdot L)$ 

Specular light:  $k_S(R \cdot E)^n$ 

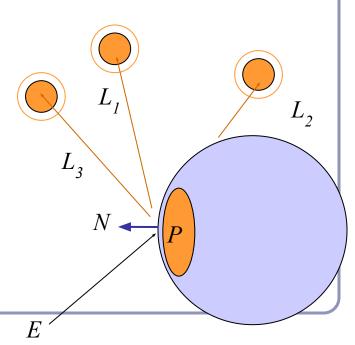
where  $R = L - 2(L \cdot N)N$ 



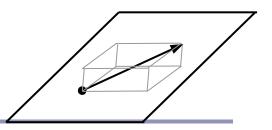
### Ray-traced illumination

Once you have the point P (the intersection of the ray with the nearest object) you'll compute how much each of the lights in the scene illuminates P.

```
\label{eq:specular} \begin{split} \textit{diffuse} &= 0 \\ \textit{for (each light $L_i$ in the scene):} \\ & \text{if } (N \bullet L) > 0: \\ & \text{[Optionally: if (a ray from P to $L_i$ can reach $L_i$):]} \\ & \textit{diffuse} += k_D(N \bullet L) \\ & \textit{specular} += k_S(R \bullet E)^n \\ & \textit{intensity at $P$ = ambient + diffuse + specular} \end{split}
```



### Hitting things with rays



A ray is defined parametrically as

$$P(t) = E + tD, \ t \ge 0 \tag{a}$$

where E is the ray's origin (our eye position) and D is the ray's direction, a unit-length vector.

We expand this equation to three dimensions, x, y and z:

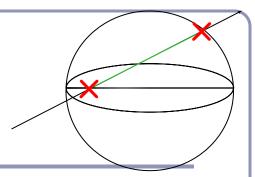
$$x(t) = x_E + tx_D$$

$$y(t) = y_E + ty_D$$

$$z(t) = z_E + tz_D$$

$$(\beta)$$

# Hitting things with rays: Sphere



The unit sphere, centered at the origin, has the implicit equation

$$x^2 + y^2 + z^2 = 1 (y)$$

Substituting equation ( $\beta$ ) into ( $\gamma$ ) gives

$$(x_E + tx_D)^2 + (y_E + ty_D)^2 + (z_E + tz_D)^2 = 1$$

which expands to

$$t^{2}(x_{D}^{2}+y_{D}^{2}+z_{D}^{2})+t(2x_{E}x_{D}^{2}+2y_{E}y_{D}^{2}+2z_{E}z_{D}^{2})+(x_{E}^{2}+y_{E}^{2}+z_{E}^{2}-1)=0$$

which is of the form

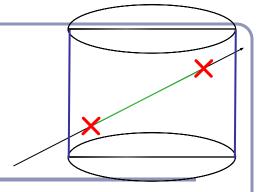
$$at^2+bt+c=0$$

which can be solved for *t*:

$$t = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a}$$

...giving us two points of intersection.

# Hitting things with rays: Cylinder



The infinite unit cylinder, centered at the origin, has the implicit equation

$$x^2 + y^2 = 1 \tag{\delta}$$

Substituting equation  $(\beta)$  into  $(\delta)$  gives

$$(x_E + tx_D)^2 + (y_E + ty_D)^2 = 1$$

which expands to

$$t^{2}(x_{D}^{2}+y_{D}^{2})+t(2x_{E}x_{D}+2y_{E}y_{D})+(x_{E}^{2}+y_{E}^{2}-1)=0$$

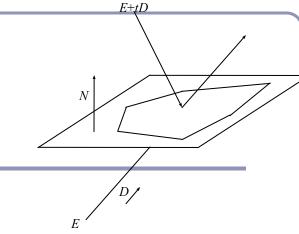
which is of the form

$$at^{2}+bt+c=0$$

which can be solved for t as before, giving us two points of intersection.

The cylinder is infinite; there is no z term.

# Hitting things with rays: Planes and polygons



A planar polygon P can be defined as

Polygon 
$$P = \{v^1, ..., v^n\}$$

which gives us the normal to P as

$$N = (v^n - v^l) \times (v^2 - v^l)$$

The equation for the plane of P is

$$N\bullet (p - v^l) = 0$$

Substituting equation ( $\alpha$ ) into ( $\zeta$ ) for p yields

$$N \bullet (E + tD - v^I) = 0$$

$$x_N(x_E + tx_D - x_v^I) + y_N(y_E + ty_D - y_v^I) + z_N(z_E + tz_D - z_v^I) = 0$$

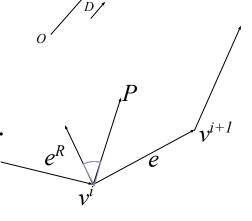
$$t = \frac{(N \bullet v^1) - (N \bullet E)}{N \bullet D}$$

 $(\zeta)$ 

### Point in convex polygon

#### Half-planes method

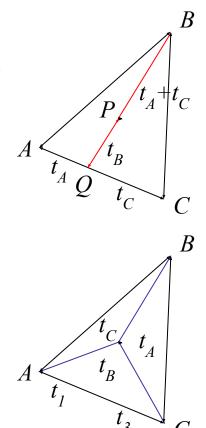
- Each edge defines an infinite half-plane covering the polygon. If the point P lies in all of the half-planes then it must be in the polygon.
- For each edge  $e=v^i \rightarrow v^{i+1}$ :
  - Rotate e by 90° CCW around N.
    - Do this quickly by crossing N with e.
  - If  $e^{R_{\bullet}}(P-v^i) < 0$  then the point is outside e.
- Fastest known method.



#### Barycentric coordinates

Barycentric coordinates  $(t_A, t_B, t_C)$  are a coordinate system for describing the location of a point P inside a triangle (A, B, C).

- You can think of  $(t_A, t_B, t_C)$  as 'masses' placed at (A, B, C) respectively so that the center of gravity of the triangle lies at P.
- $(t_A, t_B, t_C)$  are also proportional to the subtriangle areas.
  - The area of a triangle is ½ the length of the cross product of two of its sides.



### Point in nonconvex polygon

#### Winding number

- The *winding number* of a point P in a curve C is the number of times that the curve wraps around the point.
- For a simple closed curve (as any well-behaved polygon should be) this will be zero if the point is outside the curve, non-zero of it's inside.
- The winding number is the sum of the angles from  $v^i$  to P to  $v^{i+1}$ .
  - Caveat: This method is elegant but slow.

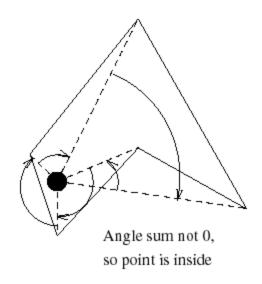
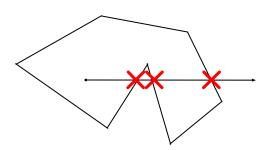


Figure from Eric Haines' "Point in Polygon Strategies", *Graphics Gems IV*, 1994

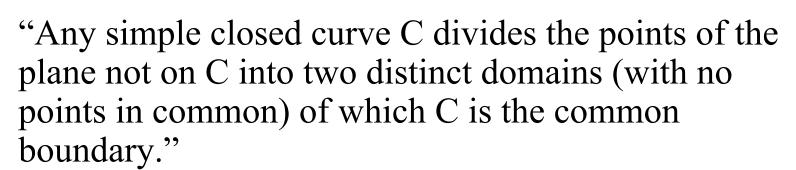
### Point in nonconvex polygon

#### Ray casting (1974)

- Odd number of crossings = inside
- Issues:
  - How to find a point that you *know* is inside?
  - What if the ray hits a vertex?
  - Best accelerated by working in 2D
    - You could transform all vertices such that the coordinate system of the polygon has normal = Z axis...
    - Or, you could observe that crossings are invariant under scaling transforms and just project along any axis by ignoring (for example) the Z component.
- Validity proved by the *Jordan curve* theorem



#### The Jordan curve theorem



• First stated (but proved incorrectly) by Camille Jordan (1838 -1922) in his *Cours d'Analyse*.

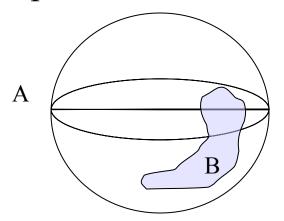
 $Sketch\ of\ proof:\ (\text{For full proof see Courant \& Robbins, 1941.})$ 

- Show that any point in A can be joined to any other point in A by a path which does not cross C, and likewise for B.
- Show that any path connecting a point in A to a point in B *must* cross C.

#### The Jordan curve theorem on a sphere

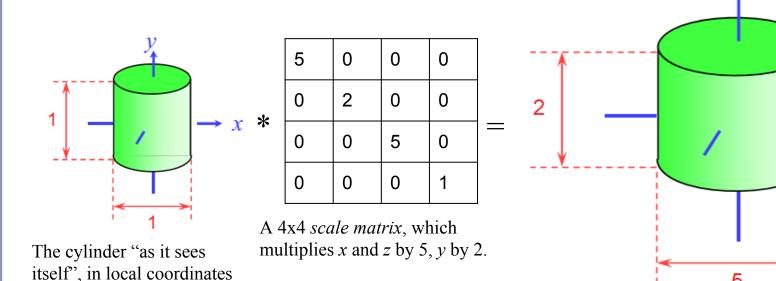
Note that the Jordan curve theorem can be extended to a curve on a sphere, or anything which is topologically equivalent to a sphere.

"Any simple closed curve on a sphere separates the surface of the sphere into two distinct regions."



### Local coordinates, world coordinates

A very common technique in graphics is to associate a *local-to-world transform*, T, with a primitive.



The cylinder "as the world sees it", in world coordinates

# Local coordinates, world coordinates: Transforming the ray

In order to test whether a ray hits a transformed object, we need to describe the ray in the object's *local* coordinates. We transform the ray by the *inverse of* the local to world matrix,  $T^{-1}$ .

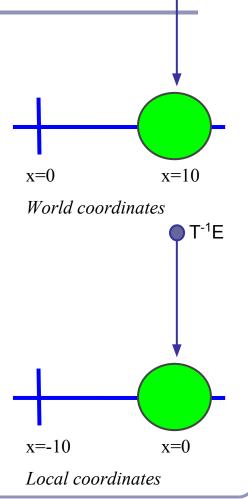
If the ray is defined by

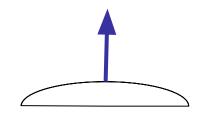
$$P(t) = E + tD$$

then the ray in local coordinates is defined by

$$T^{-1}(P(t)) = T^{-1}(E) + t(T^{-1}_{3x3}D)$$

where  $\mathbb{T}^{-1}_{3\times 3}$  is the top left 3x3 submatrix of  $\mathbb{T}^{-1}$ .





### Finding the normal

We often need to know N, the *normal to the surface* at the point where a ray hits a primitive.

• If the ray R hits the primitive P at point X then N is...

Primitive type	Equation for N
Unit Sphere centered at the origin	N = X
Infinite Unit Cylinder centered at the origin	$N = [x_X, y_X, 0]$
Infinite Double Cone centered at the origin	$N = X \times (X \times [0, 0, z_X])$
Plane with normal <i>n</i>	N = n

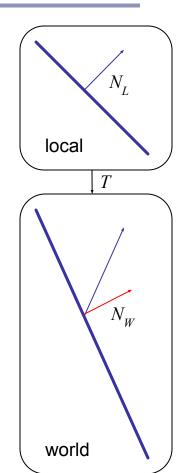
We use the normal for color, reflection, refraction, shadow rays...

# Converting the normal from local to world coordinates

To find the world-coordinates normal N from the local-coordinates  $N_L$ , multiply  $N_L$  by the transpose of the inverse of the top left-hand 3x3 submatrix of T:

$$N=((T_{3x3})^{-1})^T N_L$$

- We want the top left 3x3 to discard translations
- For any rotation Q,  $(Q^{-1})^T = Q$
- Scaling is unaffected by transpose, and a scale of (a,b,c) becomes (1/a,1/b,1/c) when inverted



## Local coordinates, world coordinates Summary

To compute the intersection of a ray R=E+tD with an object transformed by local-to-world transform T:

- 1. Compute R', the ray R in local coordinates, as  $P'(t) = T^{-1}(P(t)) = T^{-1}(E) + t(T^{-1}_{3\times3}(D))$
- 2. Perform your hit test in local coordinates.
- 3. Convert all hit points from local coordinates back to world coordinates by multiplying them by T.
- 4. Convert all hit normals from local coordinates back to world coordinates by multiplying them by ((T<sup>3×3</sup>)<sup>-1</sup>)<sup>T</sup>.

This will allow you to efficiently and quickly fire rays at arbitrarily-transformed primitive objects.

# Speed up ray-tracing with *bounding* volumes

Bounding volumes help to quickly accelerate volumetric tests, such as "does the ray hit the cow?"

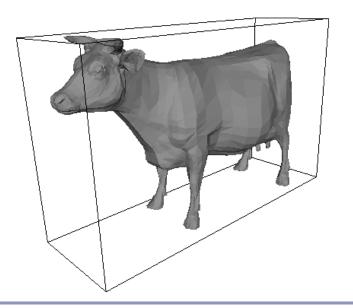
- choose fast hit testing over accuracy
- 'bboxes' don't have to be tight

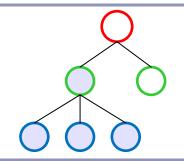
Axis-aligned bounding boxes

• max and min of x/y/z.

Bounding spheres

- max of radius from some rough center *Bounding cylinders*
- common in early FPS games

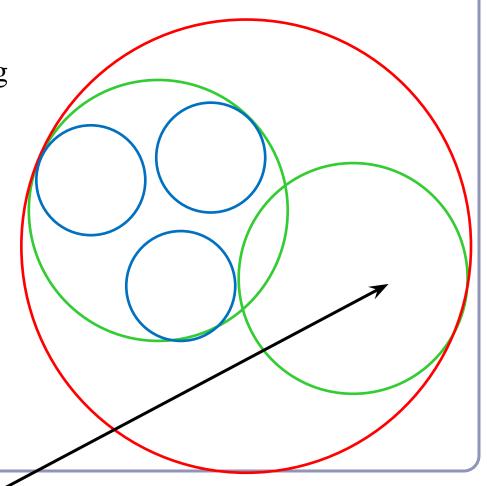




### Bounding volumes in hierarchy

Hierarchies of bounding volumes allow early discarding of rays that won't hit large parts of the scene.

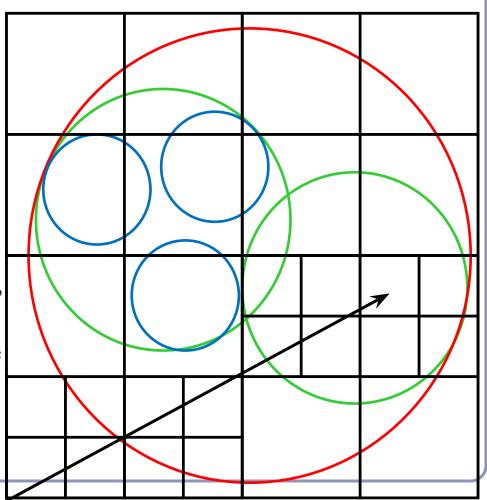
- Pro: Rays can skip subsections of the hierarchy
- Con: Without spatial coherence ordering the objects in a volume you hit, you'll still have to hit-test every object



## Subdivision of space

Split space into cells and list in each cell every object in the scene that overlaps that cell.

- Pro: The ray can skip empty cells
- Con: Depending on cell size, objects may overlap many filled cells or you may waste memory on many empty cells



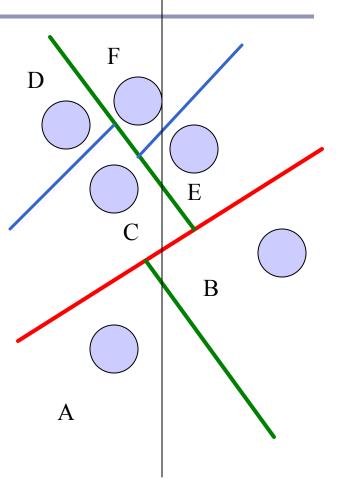
# Popular acceleration structures: BSP Trees

The *BSP tree* partitions the scene into objects in front of, on, and behind a tree of planes.

• When you fire a ray into the scene, you test all near-side objects before testing far-side objects.

#### **Problems:**

- choice of planes is not obvious
- computation is slow
- plane intersection tests are heavy on floatingpoint math.



# Popular acceleration structures: *kd-trees*

# The *kd-tree* is a simplification of the BSP Tree data structure

• Space is recursively subdivided by axisaligned planes and points on either side of each plane are separated in the tree.

The kd-tree has O(n log n) insertion time (but this is very optimizable by domain knowledge) and O(n<sup>2/3</sup>) search time.
 kd-trees don't suffer from the mathematical

• kd-trees don't suffer from the mathematical slowdowns of BSPs because their planes are always axis-aligned.

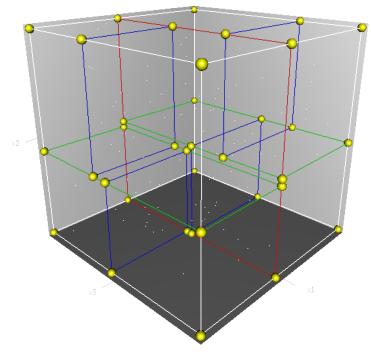


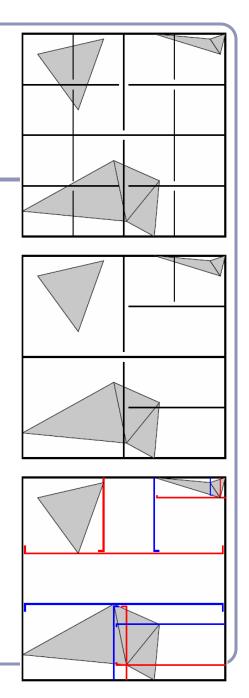
Image from Wikipedia, bless their hearts.

# Popular acceleration structures: Bounding Interval Hierarchies

The *Bounding Interval Hierarchy* subdivides space around the volumes of objects and shrinks each volume to remove unused space.

- Think of this as a "best-fit" kd-tree
- Can be built dynamically as each ray is fired into the scene

Image from Wächter and Keller's paper, Instant Ray Tracing: The Bounding Interval Hierarchy, Eurographics (2006)



## Using OpenGL to accelerate ray-tracing

To accelerate first raycast, don't raycast: use existing hardware.

- Use hardware rendering (eg OpenGL) to write to an offscreen buffer.
- Set the color of each primitive equal to a pointer to that primitive.
- Render your scene in gl with z-buffering and no lighting.
- The 'color' value at each pixel in the buffer is now a pointer to the primitive under that pixel.



#### References

#### Jordan curves

R. Courant, H. Robbins, *What is Mathematics?*, Oxford University Press, 1941 http://cgm.cs.mcgill.ca/~godfried/teaching/cg-projects/97/Octavian/compgeom.html

#### Intersection testing

http://www.realtimerendering.com/intersections.html

http://tog.acm.org/editors/erich/ptinpoly/

http://mathworld.wolfram.com/BarycentricCoordinates.html

#### Ray tracing

Foley & van Dam, Computer Graphics (1995)

Jon Genetti and Dan Gordon, *Ray Tracing With Adaptive Supersampling in Object Space*, http://www.cs.uaf.edu/~genetti/Research/Papers/GI93/GI.html (1993)

Zack Waters, "Realistic Raytracing", http://web.cs.wpi. edu/~emmanuel/courses/cs563/write\_ups/zackw/realistic\_raytracing.html